

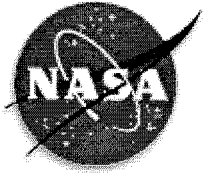
# The DRS Interferometric Displacement Sensor

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Detector Workshop, May 19-26, 2002



## Disturbance Reduction System

NMP

- Technology validation of sensor and thrust-producing technologies to control a space vehicles flight path so the payload responds only to gravitational forces.

*Sensor: Stanford University, Stanford, CA*

*Thruster: Busek Company Inc. Natick, Mass.*

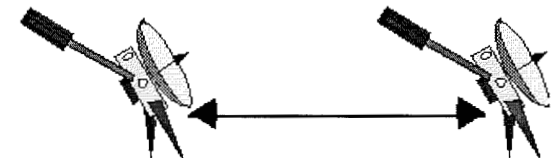
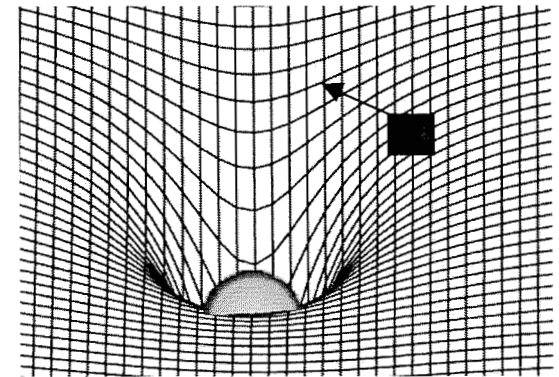
- Launch 2006 as NASA's Space Technology 7 project (ST7)
  - Piggy-backing on ESA's SMART-2 Mission



## DRS Technology Objectives

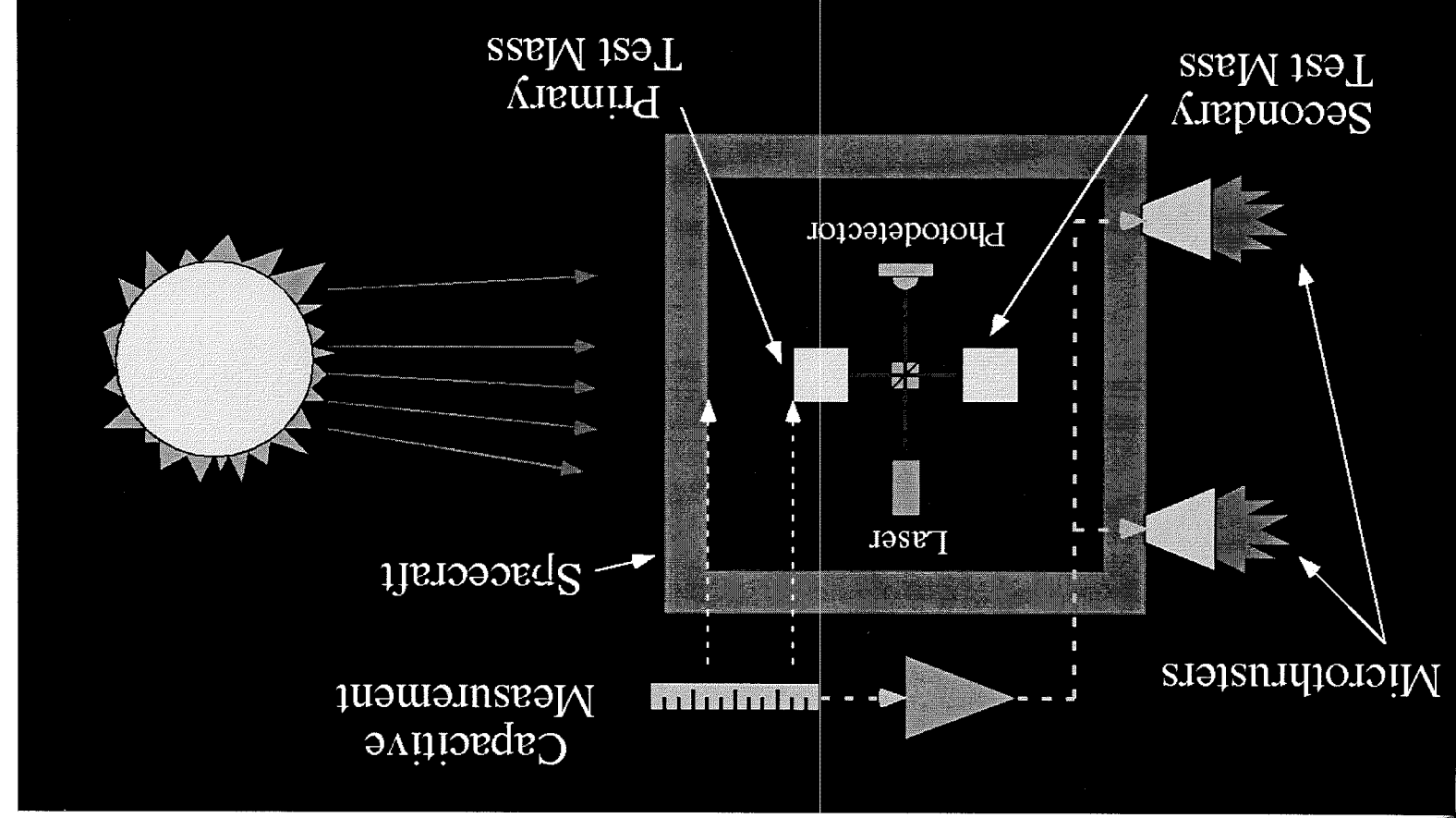


- Validate that a test mass follows a trajectory determined by gravitational forces only within  $3 \times 10^{-14} \text{ m/s}^2/\sqrt{\text{Hz}}$ 
  - Low acceleration noise is needed for study of general relativity, planetary gravity, gravitational waves
- Validate spacecraft position control to an accuracy of  $<10 \text{ nm}/\sqrt{\text{Hz}}$ 
  - Spacecraft position control is required for separated-spacecraft interferometers which do not use internal delay lines



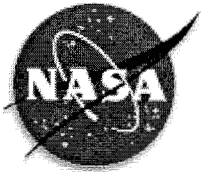


## DRS Concept



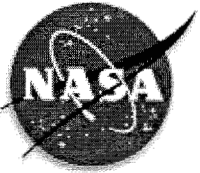
- The DRS instrument package consists of

- Two gravitational reference sensors
- Microthrusters for spacecraft position control
- Interferometer to measure the distance between the two test masses.



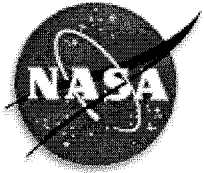
## DRS Technologies

- Gravitational Reference Sensors
  - Test mass noise  $< 3 \times 10^{-14} \text{ m/s}^2/\sqrt{\text{Hz}}$ , 1 mHz to 10 mHz
  - Measurement of position to  $< 3 \text{ nm}/\sqrt{\text{Hz}}$ , 1 mHz to 10 mHz
  - Accelerometer mode
    - Validation of thrusters
  - Force noise diagnostics
    - Validate noise models
- Micro-Newton Thrusters
  - 1-20  $\mu\text{N}$  range
  - Control precision adjustment  $< 0.1 \mu\text{N}$
  - Noise  $< 0.1 \mu\text{N}/\sqrt{\text{Hz}}$ , 1 mHz to 10 mHz
- *Interferometer is not a new technology*
  - Can be completely tested on ground
  - Is used in DRS for validation only - I.e. independent (out-of-loop) detector



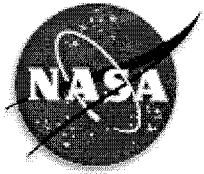
# Flight Validation Rationale

- Must be validated in space
  - 1 g environment on Earth makes ground testing impossible
- Must be in orbit far from Earth
  - Difference in gravitational force on two test masses must be small to validate instrument performance
    - Requires orbit at GEO altitude or higher
- Must have mechanically, thermally stable environment
  - Spacecraft must be in constant orientation during DRS tests
  - Thermal isolation system needed for DRS
    - Spacecraft eclipses need to be avoided
- ESA SMART-2 spacecraft suitable host
  - Will operate ~ 0.1 AU from Earth



## Flight Activities

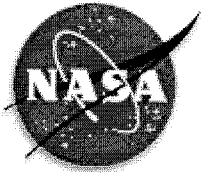
- DRS operations will take 90 days
  - DRS operates only at specific times
  - Spacecraft needs to be in quiet state
    - No maneuvers
    - No moving parts
    - No changes in power dissipation
  - Spacecraft will establish a nominal orientation and dead band
    - DRS should maintain orientation
    - If orientation dead band exceeded, DRS will be shut off
  - DRS tests will consist of a series of experiments of 1-3 days each
    - Experiments can be scheduled non-consecutively
    - Experiments can be repeated if necessary



## Flight Validation Measurements

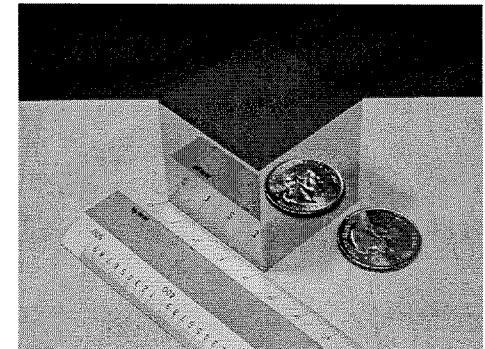
- GRS in accelerometer mode validate thruster performance
  - Can use two GRS to compare calibrations
  - Initial thruster calibration to be used for setting spacecraft control parameters
  - Calibration will be repeated several times over 6 months to check stability
- Fire thrusters to center spacecraft on one GRS test mass
  - GRS position measurements used for control and validation
  - Interferometer readout cross-checks GRS in one direction
- Fire thrusters to orient spacecraft around two test masses
  - GRS position measurements used for control and validation
- Verify acceleration noise on test masses
  - Laser interferometer compares acceleration of two test masses
- Verify force noise model verification
  - Apply known perturbations to see that response is as predicted



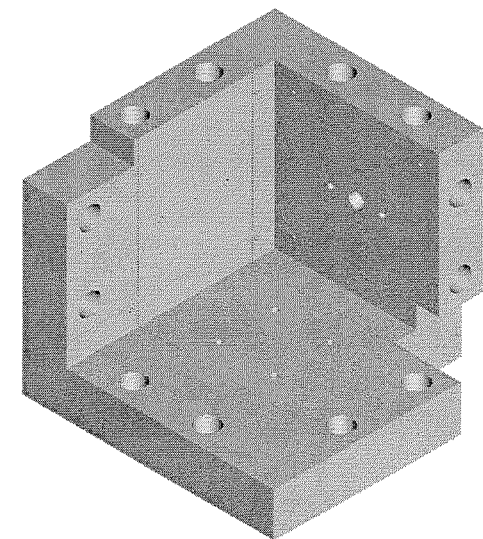


## GRS Description

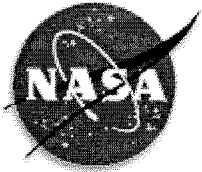
- GRS consists of;
  - A freely-floating test mass within a housing,
  - Position measurement of the test mass w.r.t. housing
  - Control of test mass orientation
  - Charge control subsystem
- The proof mass must be isolated from;
  - Solar magnetic field
  - Solar radiation pressure
  - Residual gas pressure
  - Thermal radiation pressure
  - Charging by cosmic rays
  - Spacecraft self-gravity
  - Spacecraft magnetic fields
  - Spacecraft electric fields



GRS Test Mass

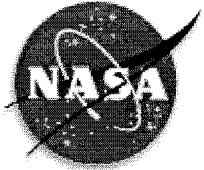


GRS Housing Assembly



## GRS Requirements

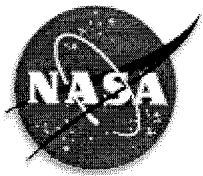
ITEM	REQUIREMENT
<b>Sensitive Axis Acceleration Noise</b>	<b><math>3 \times 10^{-14} \text{ m/s}^2/\sqrt{\text{Hz}}</math></b>
<b>Trans. Axis Acceleration Noise</b>	<b>None</b>
<b>Sensitive Axis Position Sensitivity</b>	<b><math>10 \times 10^{-9} \text{ m}/\sqrt{\text{Hz}}</math></b>
<b>Trans. Axis Position Sensitivity</b>	<b><math>2.5 \times 10^{-6} \text{ m}/\sqrt{\text{Hz}}</math></b>
<b>Orientation Control Noise</b>	<b><math>1 \times 10^{-6} \text{ rad}/\sqrt{\text{Hz}}</math></b>



## DRS Performance Limits

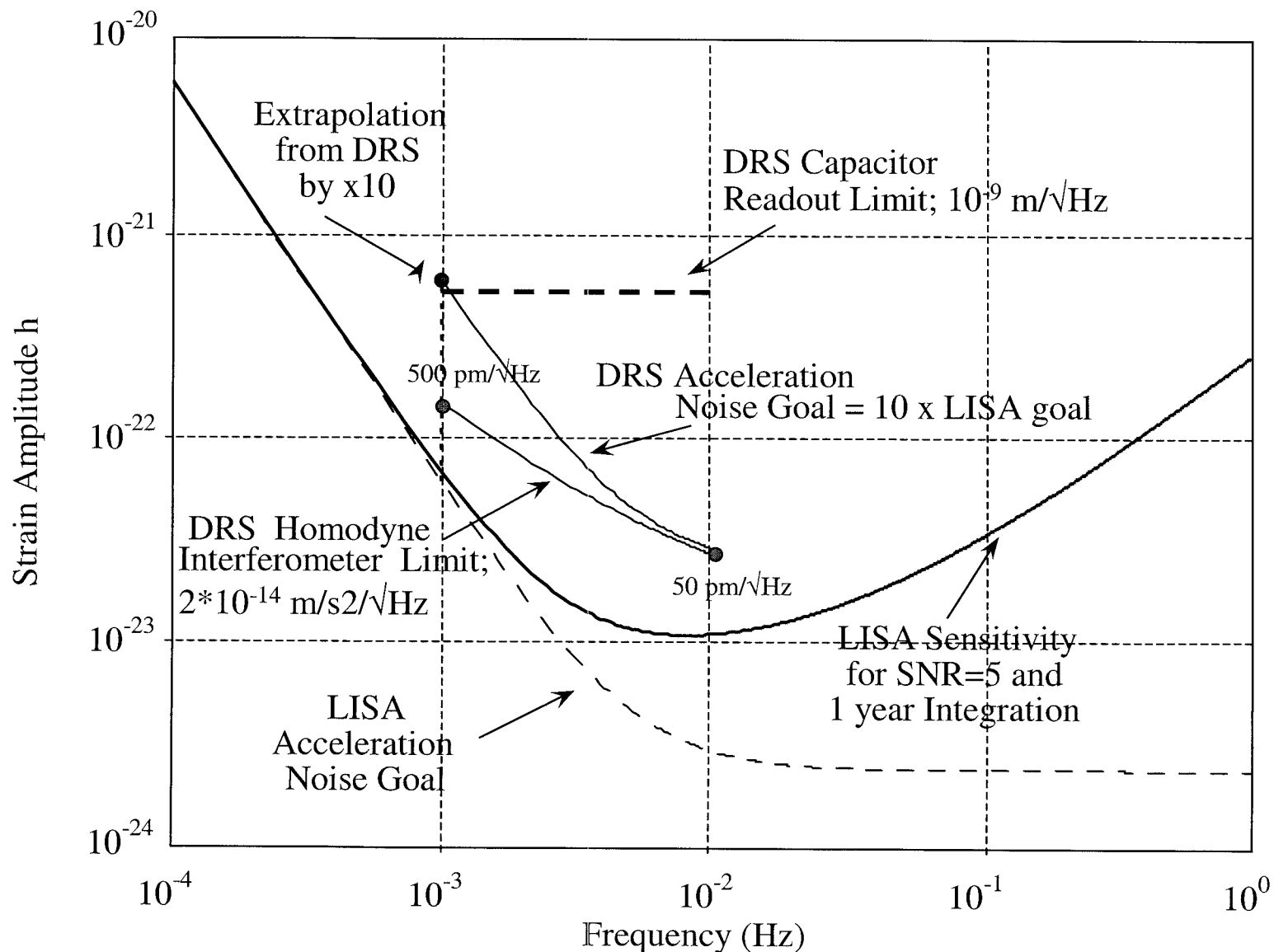
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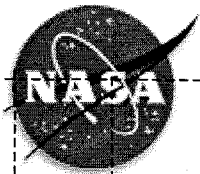
- LISA requirement @ 1 mHz:  $3 \times 10^{-15} \text{ m/s}^2/\sqrt{\text{Hz}}$
- Proof mass sensor noise:  $1 \text{ nm}/\sqrt{\text{Hz}}$
- Interferometer noise @ 1 mHz:  $500 \text{ pm}/\sqrt{\text{Hz}}$
- DRS goal @ 1 mHz:  $3 \times 10^{-14} \text{ m/s}^2/\sqrt{\text{Hz}}$
  
- LISA requirement @ 10 mHz:  $1.5 \times 10^{-14} \text{ m/s}^2/\sqrt{\text{Hz}}$
- Interferometer noise @ 10 mHz:  $50 \text{ pm}/\sqrt{\text{Hz}}$
- DRS goal @ 10 mHz:  $1.5 \times 10^{-13} \text{ m/s}^2/\sqrt{\text{Hz}}$
  
- $a = (2\pi f)^2 \cdot 5 \times 10^{-10} \text{ m}/\sqrt{\text{Hz}} \sim 2 \times 10^{-14} \text{ m/s}^2/\sqrt{\text{Hz}}$  @ 1 mHz
- $a = (2\pi f)^2 \cdot 1 \times 10^{-9} \text{ m}/\sqrt{\text{Hz}} \sim 4 \times 10^{-14} \text{ m/s}^2/\sqrt{\text{Hz}}$  @ 1 mHz
- $4 \times 10^{-12} \text{ m/s}^2/\sqrt{\text{Hz}}$  @ 10 mHz)
- $a = (2\pi f)^2 \cdot 5 \times 10^{-11} \text{ m}/\sqrt{\text{Hz}} \sim 2 \times 10^{-13} \text{ m/s}^2/\sqrt{\text{Hz}}$  @ 10 mHz



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# DRS Performance Limits



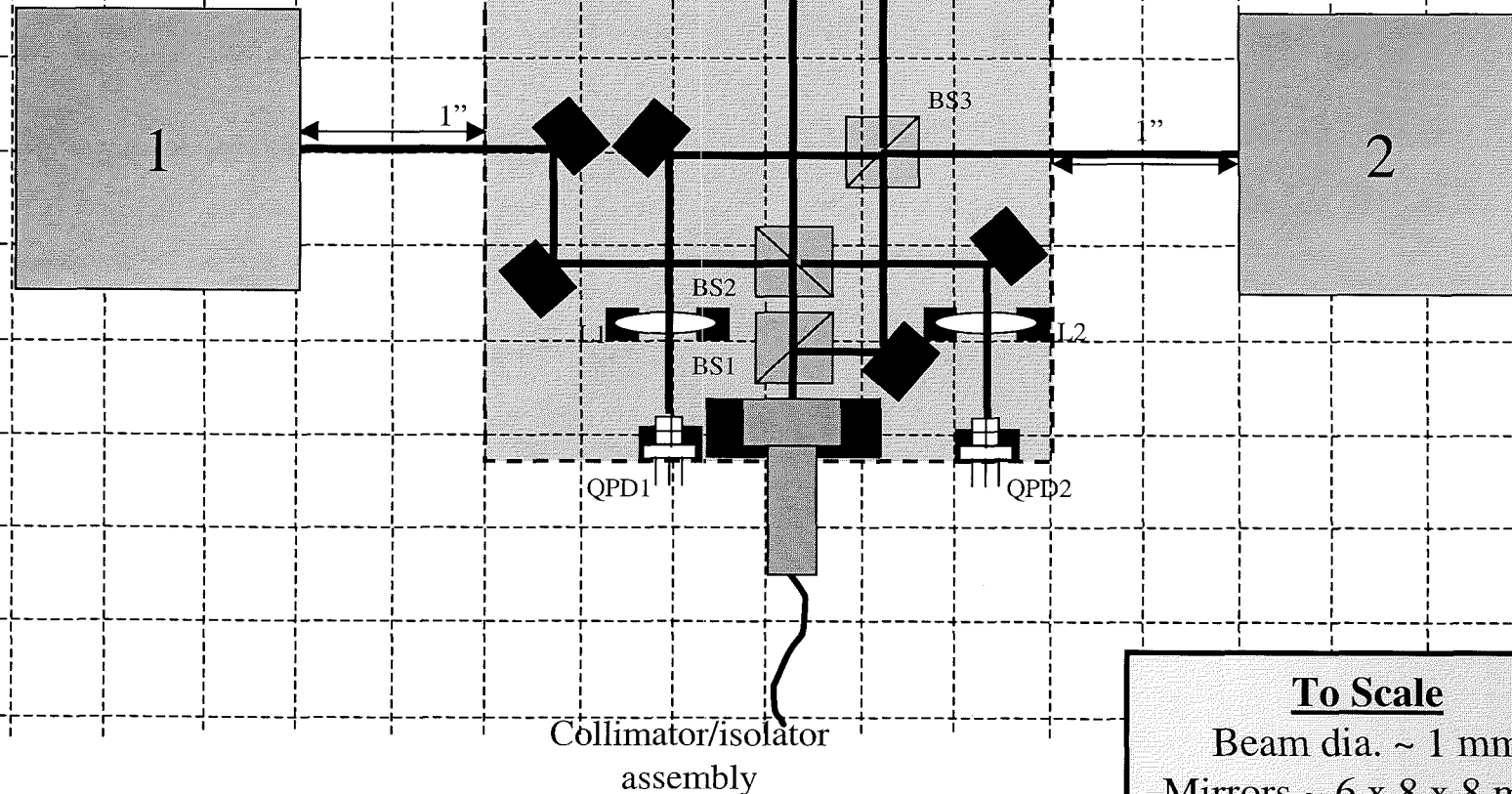


# Interferometer Baseline Design

0.5"

0.5"

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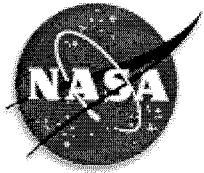


## To Scale

Beam dia. ~ 1 mm

Mirrors ~ 6 x 8 x 8 mm<sup>3</sup>

BS cubes ~ 8 x 8 x 8 mm<sup>3</sup>



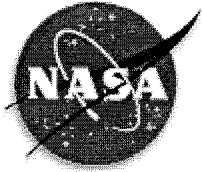
## Laser Requirements

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- Average unstabilized laser noise seen:  $\sim 5.5 \text{ MHz}/\sqrt{\text{Hz}}$  @ 10 mHz
- $\sim 30 \text{ MHz}/\sqrt{\text{Hz}}$  @ 1 mHz
- Interferometer pathlength mismatch:  $\Delta L < 1 \text{ mm}$
- Laser noise  $\rightarrow$  path noise:  
@ 10 mHz  $\Delta x = \Delta L * \Delta v / v$   
@ 1 mHz  $< 20 \text{ pm}/\sqrt{\text{Hz}}$  (avg.)  
 $< 100 \text{ pm}/\sqrt{\text{Hz}}$  (avg.)

Compare to:

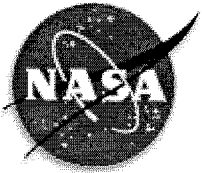
- Interferometer noise req'd:  $50 \text{ pm}/\sqrt{\text{Hz}}$  @ 10 mHz
- Interferometer noise req'd:  $500 \text{ pm}/\sqrt{\text{Hz}}$  @ 1 mHz



## Proof Mass Stability

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- Proof mass pointing stability:  $1 \mu\text{rad}/\sqrt{\text{Hz}}$  @ 1 mHz
  - associated displacement error: TBD



# OPD error vs. Fringe Visibility/wavefront tilt

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- OPD error due to changing fringe contrast @ max. slope of fringe signal (e.g due to proof mass tip/tilt):

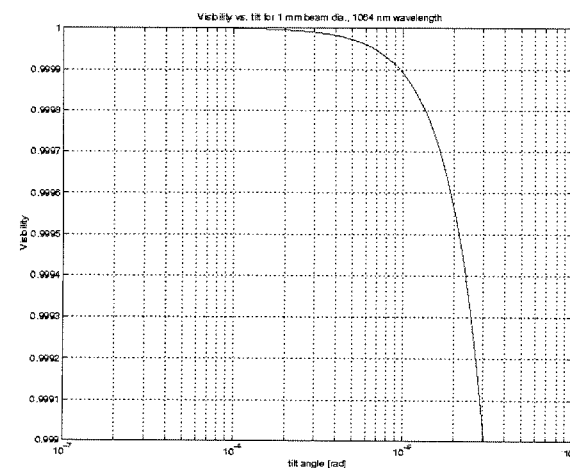
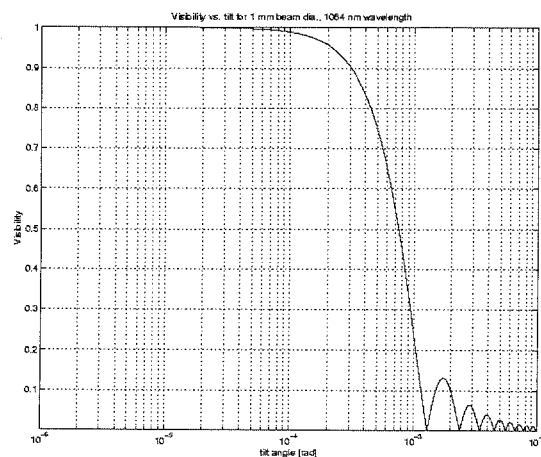
$$\text{Fringe signal: } I = I_0 * (1 + V * \sin(4\pi * x / \lambda))$$

$$dI/I_0 = 4\pi V / \lambda * \cos(4\pi * x / \lambda) dx + \sin(4\pi * x / \lambda) dV$$

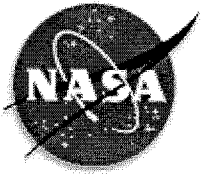
=> @  $x = n * \lambda / 4$  w/  $n = 1, 2, 3, \dots$  (the max. slope of the fringe)  $\sin(4\pi * x / \lambda) = 0$ ,

i.e. no sensitivity to visibility change  $dV$ . Q: How close can we get to that?

- With  $V^2 = [2 * J_1(\pi * \theta * D / \lambda) / (\pi * \theta * D / \lambda)]^2$  w/  $\theta$  angle between wave fronts,  $D = 1$  mm,  $\lambda = 1064$  nm we get a  $dV \sim 10^{-5}$  (10 ppm) with the  $1 \mu\text{rad}/\sqrt{\text{Hz}}$  @ 1 mHz proof mass pointing stability.







## OPD error vs. Signal (laser) Intensity

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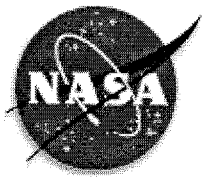
- OPD error due to changing laser intensity  $I_0$  :

$$\frac{dI}{dI_0} = 1 + \sin(4\pi * x / \lambda) \text{ w/ } \sin(4\pi * x / \lambda) = 0$$
$$@ x = n * \lambda / 4$$

$$dI = dI_0$$

$$\text{using } dI/I_0 = 4\pi V / \lambda * \cos(4\pi * x / \lambda) * dx$$

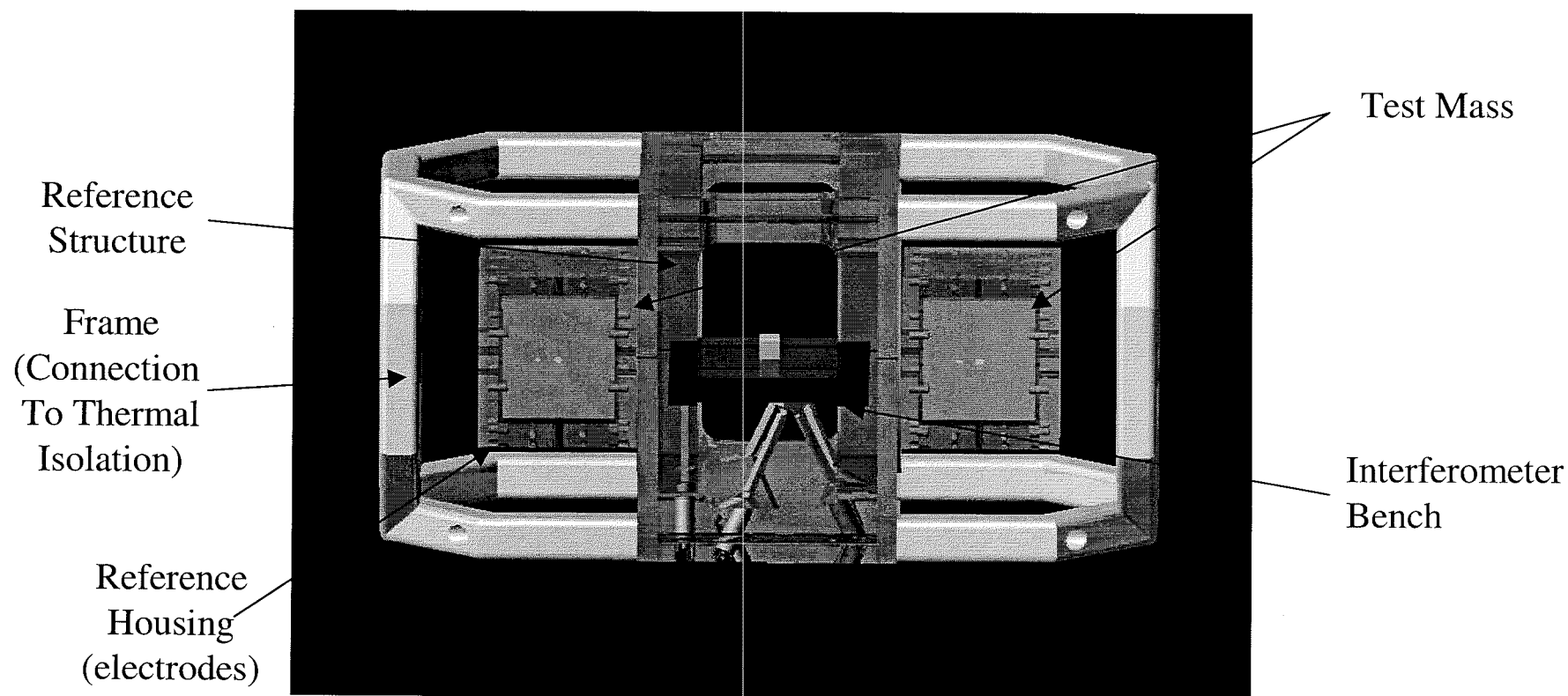
we get an error of  $dx = V * 20 \text{ pm}$  for a signal (laser) intensity change  $dI_0 \sim 2 * 10^{-4}$  @  $x \sim n * \lambda / 4$ .

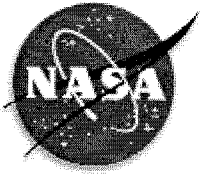


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## Instrument Configuration

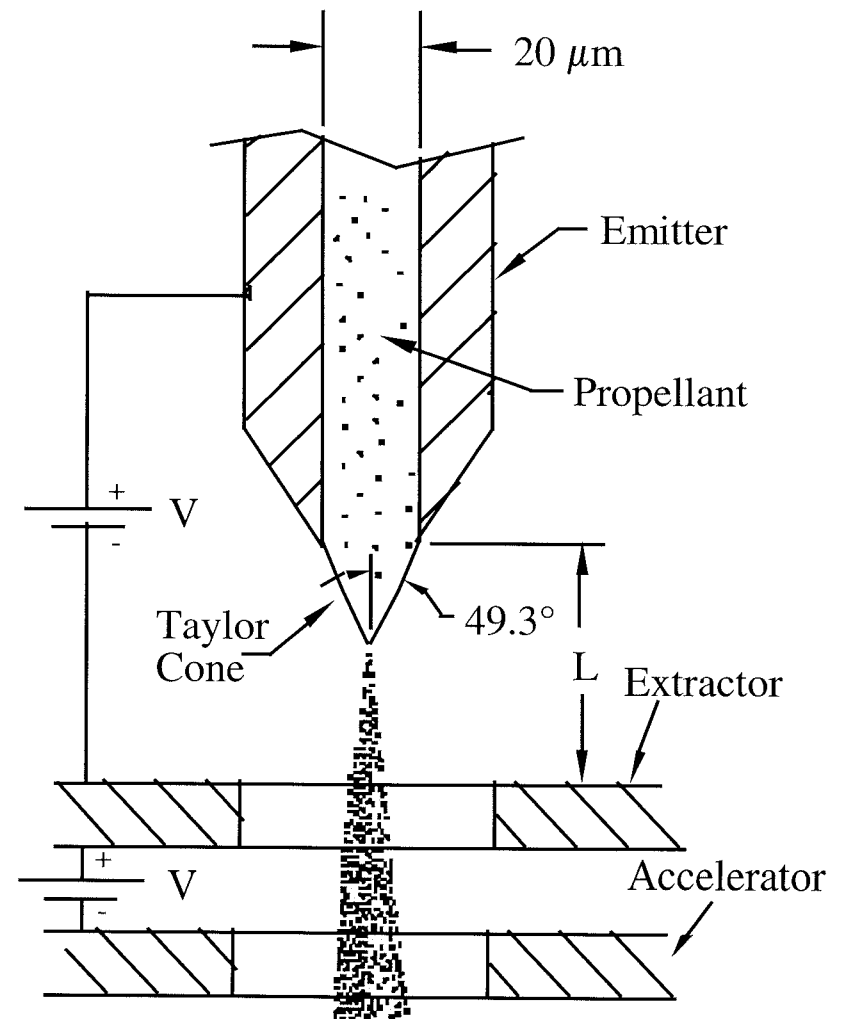
- Two gravitational reference sensors.
- Individual vacuum housings (not shown).
- Sensors integrated with interferometer.

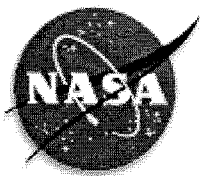




# Microthruster Concept

- Colloidal thrusters
  - Fluid fed through fine needle
  - High voltage extracts charged droplets
  - Droplets accelerated by high voltage
  - Thrust control of  $0.1 \mu\text{N}$  through change of voltage or flow rate
  - Many needles can be combined to develop necessary thrust
  - Proportional thrust control allows precision position control
    - (displacement to thrust feedback loop)





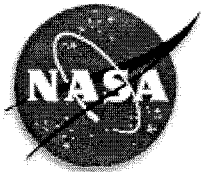
# Microthruster Requirements

#	Item	Specification	Comment
1	Thrust	1– 20 $\mu$ N	Smoothly variable between end point values within 0.1 $\mu$ N
2	Thrust controllability/resolution	0.1 $\mu$ N	Must be within $\pm 0.1\mu$ N from set point
3	Thrust noise	0.1 $\mu$ N/ $\sqrt{\text{Hz}}$ from $10^{-3}$ Hz to $10^{-2}$ Hz	Stable over given period
4	Specific impulse	$\sim 500$ sec	May vary depending on thrust
5	Thrust slew rate	$< 0.5 \mu\text{N/sec}$	Over voltage range/slower with flow
6	System mass	$\leq 2$ kg	Includes all thruster subsystem
7	Total system power	$< 6.2$ W	Average/thruster zeolite heater major consumer
8	Total operating time	$\geq 2000$ hours	90 days mission
<b>Propulsion System Package</b>			
<ul style="list-style-type: none"><li>– 8 thrusters in 4 clusters</li><li>– Continuous/differential operation, authority over 6 DOF</li></ul>			

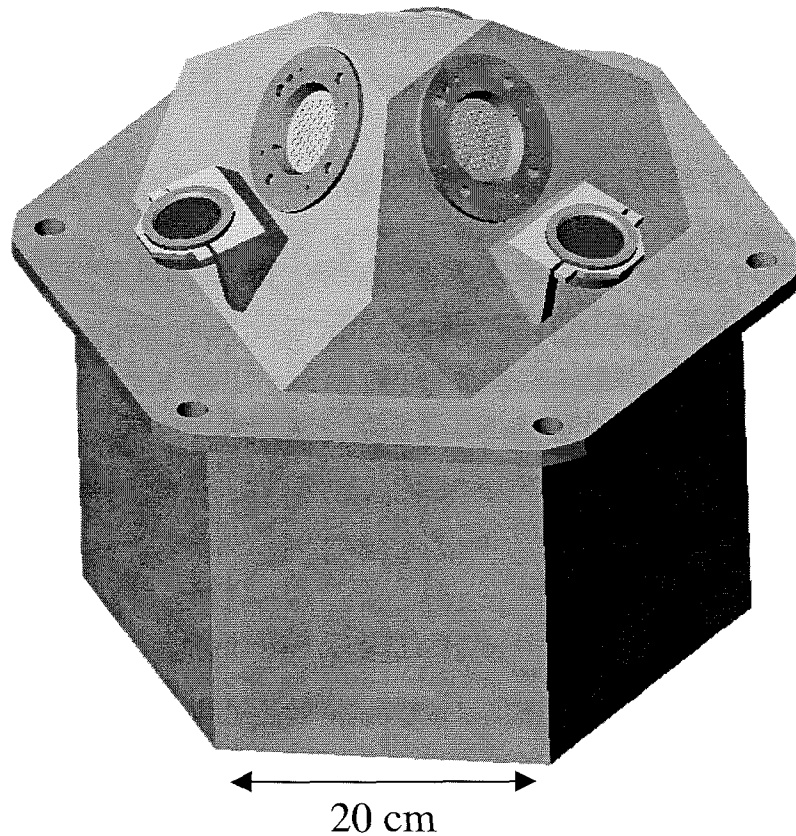


## Microthruster Ground Test

- Performance can be characterized in ground tests
  - Thrust levels inferred from beam current (time of flight spectrometry)
  - Torsional thrust stand verifies thrust to sub- $\mu\text{N}$  levels (also at JPL)
  - Lifetime tests (emitter/electrochemistry and extractor/sputtering)
  - Mission profile simulation tests, life and dynamic response demonstration (thrust commands from simulated spacecraft computer)
  - Beam neutralization and beam profile measurements
    - Input to models to predict behavior in space
  - Contamination tests
    - Limited by back-scatter from test chambers (JPL)



## Microthruster Flight/Validation Experiments



3 Thrusters Cluster concept

2 Thrusters Cluster baselined

- Upon final orbit acquisition place GRS's in accelerometer mode
- Fire one thruster at a time
- Sweep 1 to 20  $\mu\text{N}$ , compute calibration factors
- Hold steady  $t \geq 1000$  sec. and record thrust noise at 3 levels of thrust
- Validate ground tests (thrust, noise, beam/neutralization)
- Repeat at middle and end of DRS operations